# The Reflex-Diode HPM Source on Aurora

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Abstract—This paper describes the most recent in a series of experiments to develop the reflex diode as a source of microwaves on the Aurora relativistic electron-beam pulser. We have achieved an overall output for radial extraction of  $\sim 400$  J in microwave bursts from  $\sim 100$  to 150 ns at frequencies below 1 GHz. The diagnostics for radial extraction have included directional couplers, card calorimeters, and free-field sensors. We have varied the anode/cathode spacing, downstream microwave reflector, and a second anode foil, but, within the range of variations, no strong trends have been noted.

# I. INTRODUCTION

A URORA is one of the largest pulsed, high-intensity electron-beam generators. The machine consists of four parallel Blumlein pulsers, each producing a 180-ns (FWHM) electron beam peaking at about 10 MeV and 250 kA with a  $\sim 35-\Omega$  load. Although Aurora usually produces bremsstrahlung, it has been interesting to investigate alternative applications, one of which is microwaves. Because of the size of hardware compatible with the 1.22-m (diam) magnetically insulated transmission line (MITL), we have restricted our attention to the reflex diode, which is the simplest mechanism for microwave production. Despite the inefficiency of this mechanism, the magnitude of our power source still suggested we would produce substantial levels of microwave radiation.

With the generic reflex-diode oscillator [1] the beam is shot through a thin anode into an evacuated drift tube. A virtual cathode forms in the drift tube if the current exceeds the space-charge limit, as expected from simple theory [2] backed by particle-in-cell (PIC) simulations. The presence of a virtual cathode results in electron reflexing and microwave generation. The simulations also show that the virtual cathode itself can oscillate, giving another mechanism for microwave generation.

The initial experiment [3] on Aurora, conducted in August 1986, featured a 5-m-long axial wire calorimeter, which showed approximately 1 kJ of microwaves at the 20-GW power level. The high background in the axial

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calorimeter measurements prompted a second experiment [4] in which the axial calorimeter viewed microwaves reflected from a 45° screen in a 1.22-m diam "T" section, while the beam and debris were to have passed through the screen. Unfortunately, the axial calorimeter showed only  $\sim 50$  J of microwave energy. We believe that the lack of a sufficient drift space beyond the reflection screen for the beam and debris was partially responsible for this low result. Each of these two experiments was supplemented by PIC calculations [5], which generally agreed with the measured diode impedance, microwave frequency, microwave turn-on, and the trend of microwave energy with the foil thickness.

We conducted a third experimental session [6] on Aurora in June 1988. Because of its difficulty we abandoned axial calorimetry in favor of more definitive measurements of the microwave energy that could be extracted by radial rectangular waveguides. The microwave power and energy per arm was seen to be independent of the number of arms up to the eight that were available. We tried using a carbonized-felt cathode coating and found no improvement in microwave output. We also tried different anode materials, of which we had the most success with 16-µm Al foil instead of the more common screens. We have subsequently found that a 16- $\mu$ m foil can survive tens of shots without beam damage if we divert to a dummy load the energy from the trailing edge of the electrical pulse. This longevity of the anode foils has made it possible to keep vacuum cycling to a minimum and to shoot at a reasonably high rate.

The fourth experiment, conducted in March 1989, is the subject of this paper. Our goal in this experiment was to verify levels of radiated power and energy density and to study some of the effects of simple variations in geometry.

# II. EXPERIMENTAL DESCRIPTION

The general layout of the most recent experiment as shown in Fig. 1 was similar to the previous Aurora highpower microwave (HPM) experiment [6]. We replaced one of the four bremsstrahlung diodes with a 45° elbow plus an additional 2 m of ~50- $\Omega$  MITL leading to the uncoated cathode tip. The elbow was necessary to align the MITL with the floor, and the 2-m section downstream of the elbow was assumed to be sufficiently long to isolate the beam from the asymmetries of the elbow. The 53-cm-

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Fig. 1. Experimental configuration. (a) Plan view. (b) Axial view.

diam cathode with a 2.5-cm beveled tip and 30-cm-diam hole on axis was the same one used in the initial work. The anode foil covered the entire 1.22-m diam of the drift tube except along the edge where an array of semicircular pump holes were cut.

The anode/cathode gap was typically 23 cm. On occasion, we placed an additional foil at positions from 9 to 76-cm downstream from the anode. This foil was intended either as a mechanism to extend the reflexing path length [7], [8] or as a microwave reflector for resonance. The specific function of the secondary foil depended on the foil's position relative to the virtual cathode, assumed to be  $\sim 12$  to 20-cm downstream from the anode foil as indicated by PIC simulations using the code MAGIC [9]. Otherwise, we permitted the microwaves to propagate axially for several meters.

For microwave extraction we surrounded the axial position 27-cm downstream from the anode with six WR-975 rectangular waveguides (605-MHz cutoff) oriented with  $E_{TE_{10}}$  parallel to the z-axis of the drift chamber. Some of the six waveguides included 60° *H*-plane elbows, and one waveguide included a 90° *E*-plane bend. We also included an additional pair of diametrically opposed rectangular waveguides 3-m downstream from the anode, raising to eight the total number of waveguides loading the microwave source.

Except where the energy radiated through open-ended waveguides, resistive cards [10] absorbed the waveguide microwave energy. In our previous experiment [6] the cards had been parallel to and located at the ideal 850-MHz  $\lambda/4$  distance from the end plate, but network analysis had shown rather high reflections  $S_{11}$  ranging from

-8 to -10 dB between 0.7 and 1.0 GHz. For the present experiment, we empirically reduced  $S_{11}$  below -15 dB in the  $\sim$  700- to 950-MHz range by moving the cards to the nominal  $\lambda/8$  position. We attempted to sense the temperature change in each card through a pair of series thermistors connected to a Wheatstone bridge in the shield room. These cards were well-behaved when calibrated with tens of joules of microwaves at  $\sim 50$  W and seem to have worked well for the previous experiment [6]. However, in the present experiment the calorimetry was sufficiently erratic so that it obscured trends in the data. The possibility of high-power breakdown initiated by the thermistors and their leads is a concern and might be more critical in the new position of the cards. Although we tried to keep the leads perpendicular to  $E_{TE_{10}}$ , it is possible that higher-order modes might have coupled directly to the thermistor circuit. We have begun work on a new calorimeter designed to rectify these problems. In general, the calorimeters did show energy levels within a factor of two of the levels obtained from the directional couplers and they did respond properly to being screened from microwaves.

Vestigial-loop directional couplers  $\sim 1$ -m upstream of each card termination detected microwaves flowing away from the source. Each of these couplers is a flattened loop recessed in a cylindrical cavity (below cutoff) in the center of the H-plane side of the waveguide. The loop was designed to have a response of equal magnitude to the electric and magnetic fields so that, at a given end of the loop, there would be no signal for a particular direction of propagation in the waveguide. Therefore the signal at one end of the loop would correspond to microwave propagation in one direction, while the signal at the other end would correspond to propagation in the other direction. The directivity was better than 15 dB for frequencies above 650 MHz, and better than 30 dB for frequencies above 750 MHz. For the most recent work we dummyloaded the side of the loop corresponding to microwaves returning to the source. High-power reflection data from the cards used in the first two Aurora microwave experiments are commensurate with the range of reflections expected from the network analysis. Therefore we did not dedicate any data channels to measure reflected microwaves in the more recent experiments.

The signals from the directional coupler propagated to the shield room through 120 ns of ~ 1-cm foam-insulated cable. To keep down the radiation effects, we shielded the cables with lead "chimneys" and steel-covered trenches. Also, 300-MHz high-pass filters eliminated lowfrequency noise, and, in most cases, 1-GHz filters restricted the data to the TE<sub>10</sub> mode in the WR-975 waveguide. We monitored the directional couplers with diode detectors and transient digitizers<sup>1</sup> on all shots, and, as other instrumentation needs permitted, we also used 1-GHz oscilloscopes.<sup>2</sup> We digitized the amplitudes of the

<sup>1</sup>Tektronix 7912. <sup>2</sup>Tektronix 7104. oscilloscope traces to verify the power and energy given by the diodes. We used 500-MHz low-pass filters on the diode outputs to eliminate the RF component. For on-line Fourier analysis we mixed the signals with a local oscillator at 600 MHz (just below the waveguide cutoff) and recorded these with transient digitizers.

We calibrated the directional couplers and their data channels with a vector network analyzer.<sup>3</sup> Since some stages of the analysis involved the measurement of 80- to 90-dB attenuations, we checked the network analyzer against comparable, well-calibrated broad-band attenuators and found that, at worst, the network analysis introduced an estimated error of  $\pm 0.2$  dB. The overall responses of the directional-coupler channels up to and through the oscilloscopes are shown in Fig. 2(a), while the responses up to the diodes are shown in Fig. 2(b). We measured these responses after the final shot of the test sequence using the complete lines with the devices as attached for the last shot. (Only couplers 1, 2, and 6 were routed to the1-GHz oscilloscopes during the experiment.) The dips in Fig. 2(a) at 1.1 GHz resulted from the connection of 1.5- and 1.0-GHz filters by the 6-dB splitters that routed the signals to the oscilloscopes. The deviation of the high-frequency trend of the coupler 2 curve from the others was caused by a difference in the WR-975 setup during the calibration. The high-frequency difference between the diode-channel calibration curve for coupler 4 and the others was caused by a different type of 1-GHz filter.

To measure the radiation from either one or two openended waveguides, we placed a line of 10 monopole antennas in a  $3 \times 3 \times 6$ -m<sup>3</sup> anechoic chamber as indicated in Fig. 1(b). The anechoic chamber was a temporary structure lined with conical absorbers. These absorbers were up to 60 cm in length, which is less than optimum. However, since our anechoic chamber was intended as an absorber of radiated energy and not as a simulation of unbounded space, we believe this compromise posed no serious problem.

Each monopole antenna consisted of an RG-141 center conductor protruding 1 mm above a  $30 \times 30$ -cm<sup>2</sup> metallic ground plane and oriented parallel to the electric field vector. We calibrated each antenna on-axis 3.66 m in front of an open-ended guide using the network analyzer in conjuction with a 50-W, 0.1- to 1.0-GHz amplifier. This distance, 3.66 m, was the same as used between the openended guides and the line of receiving antennas for the HPM experiments. This calibration made the antennas reasonably accurate sensors of radiated power and energy, even though the chamber was less than ideal. However, since the fields at the antennas depended on wall reflections and our antennas were not calibrated directly in terms of known fields, the usage of the antennas for field measurements was less accurate. For the power densities quoted in this paper we have assumed that the anechoic chamber was ideal. Frequency responses of two of

<sup>3</sup>Hewlett-Packard Model 8510 A.



Fig. 2. Forward responses of vestigial-loop waveguide directional couplers: (a) oscilloscope channels (WR-975 to oscilloscope display); and (b) diode-detector channels (WR-975 to diodes).

the monopole antennas to total radiated power are shown in Fig. 3. The antenna signals were recorded directly on 1-GHz oscilloscopes.

Fig. 4 shows the electron-beam characteristics as measured with a self-integrating capacitive voltage divider and current-viewing resistor (CVR) just upstream of the elbow. We inferred the diode characteristics from these data using a transmission-line circuit analysis, which is a slight modification of the usual *Li* correction. The fast decay of the trailing edge of the voltage pulse resulted from the energy-diverting switch at the output of the pulse-forming network (PFN). The diode voltage peaked at -6.5 MV and the diode current peaked at -0.3 MA, implying a diode impedance of 22  $\Omega$ . We obtained the instrumentation trigger and fiducials from the CVR just upstream of the elbow.

Fig. 1(a) also shows a set of three electric-field sensors. However, as the data from these are suspect, we will replace the sensors with *B*-dot loops for future experiments. We terminated the axial drift region with ferrite tiles<sup>4</sup> to keep the axial microwave propagation unidirectional, which was to have facilitated the interpretation of the data from the electric-field sensors in terms of microwaves. However, we are concerned that these tiles might have saturated in the microwave flux.

<sup>4</sup>Emerson & Cuming Eccosorb<sup>®</sup> NZ-31.



#### ig. 4. Typical fellex diode characteristics (shot 0842

# III. RESULTS/DISCUSSION

Our data from March 1989 comprise a total of 68 shots and include several independent measurements of the total power and energy extracted by the waveguides. The shots described below typify the data.

For shot 6840 we used only one foil, and, as for two other shots, only the lower (arm 5) of the two radiating arms was open to the anechoic chamber. Also, the signals were filtered below the 1.2-GHz cutoff of the  $TE_{01}$  and  $TE_{20}$  modes in the arm. Therefore, other than for an echo from the less-than-ideal "anechoic" walls, there was no possibility for a time-dependent distribution of power density in the chamber.

This simplicity of the radiated field made it possible to use one of the ten monopole antennas to infer the total radiated waveguide power and energy. Antenna 3 was almost directly aligned with the radiating waveguide and at the same distance from the waveguide as for its calibration. Therefore we feel confident that the data from antenna 3 are a reliable indicator of the total radiated power. Antenna 4 was 0.3-m out of alignment with the radiating arm. Therefore we made a small correction using the ideal free-space pattern for an open-end waveguide so that it too would give the total power radiated. The envelopes of the signals from antennas 3 and 4 are shown in Fig. 5(a). The frequency from the original signals was  $750 \pm 60$ MHz. The two measurements of the waveguide power have a similar time-dependence and agree with an amplitude of ~ 1.0 GW. Any difference resulted from the rather strong frequency dependence of the antenna calibrations and our inability to determine the frequency with better accuracy.

For comparison, Fig. 5(b) shows the same power envelope measured with the directional coupler in the radiating arm. Both the shape and amplitude of power agree with the antennas. We had been concerned with the possibility of breakdown in the waveguide directional couplers at the circular apertures to the vestigal loops. These apertures are rounded, but include an insulating cover which might flash over in field components parallel to the surface. This breakdown might be a distinct possibility with high-frequency overmoding, which might also cause problems with an abnormally strong coupling to the loop and terminator. We are encouraged, however, by the agreement between the directional coupler in arm 5 and the free-field antennas and by the general agreement in the past between the couplers and calorimeters.

Fig. 5(c) shows the running integrals of the power from the two antennas and directional coupler. These different measurements agree with a final energy of about 60 J.

The data for each of the other arms on shot 6840 are shown in Fig. 6. In Fig. 6(a) the directional coupler data show somewhat differing power profiles. Since arm 6 (going to the upper open-ended waveguide) was screened off at its source end, its trace illustrates either the level of noise in the diagnostics or a baseline drift. No trace is given for coupler 3 because the transmission line from that coupler was malfunctioning on that shot. Fig. 6(b)shows the energy in the various arms from the directional couplers. The average energy per arm is seen to have been about 55 J; subtracting the background suggested by the trace for arm 6 gives the total energy extracted in the four recorded arms as being about 200 J.

The total energy as given by the free-field antennas and directional couplers was a little higher than given by our previous data [6]. This might have resulted from the longer duration of the microwaves, which more recently approached the nominal width of the voltage and current. One improvement in preparation for the most recent experiment had been rectangular-waveguide apertures smoothed to  $\sim 2$ -mm radii. Assuming the static field configuration, this corresponds to a field enhancement factor of  $\sim 7$ . Previously, the apertures had been rather sharp, with larger and uncertain enhancement; however, the decreased enhancement might account for the longer duration of the radially extracted microwaves.

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Fig. 5. Comparison of antenna and directional-coupler measurements (Shot 6840). (a) Radiated power sensed by each of the two antennas most closely aligned with the single radiating arm. (b) Directional-coupler measurement of radiated power. (c) Energy radiated from each of the antennas and directional coupler.

The 0.13-MJ diode energy (as summed up to the time the PFN output was diverted) indicates that the four waveguide arms illustrated in Fig. 5 had a total efficiency of 0.15%. If we assume that the arms not monitored by directional couplers had similar energies, then we can quote an efficiency approaching 0.3%. Note that this efficiency only includes radially extracted microwaves between 0.6 and 1.0 GHz and assumes that our traces include the entire microwave signals.

While the field map with only one radiating arm was relatively stable as expected, such was not the case with two arms. The power densities recorded by the two adjacent antennas 3 and 4 on shot 6838 (Fig. 7) show a time-



Fig. 6. Comparison of directional-coupler measurements (shot 6840): (a) forward power; and (b) energy.



Fig. 7. Power density as recorded by two adjacent receiving antennas in the field of two radiating arms (shot 6838).

dependent radiation pattern. The temporal discrepancy of the data from the two antennas is evidence of the changing phase difference between the two radiating arms. We would have expected this phase difference to change from



Fig. 8. Sample frequency determination on arm 2 for shot 6872: (a) Heterodyned signal as seen on digitizer; and (b) resulting spectral energy density (normalized to give the energy detected by diode).

the lack of correspondence between the different profiles of the arms in Fig. 6(a) and from the arm-to-arm phase difference changes seen in our past work [6]. The energy densities recorded at antennas 3 and 4 were 0.67 and 0.54  $(\pm 0.15)$  J-m<sup>-2</sup>, respectively.

Shot 6872 produced one of our cleanest spectra, obtained from the heterodyned signal shown in Fig. 8(a) recorded on the transient digitizers sweeping at 5 ns/div. Because of this sweep speed, unfortunately, the complete intermediate-frequency signal could not be recorded. This signal resulted in the spectrum shown in Fig. 8(b). Since we had no calibration information for the mixers, we weighted the square of the Fourier transform so that its integral agreed with the integrated power from the diode detector.

Prompted by previous work at the Lawrence Livermore National Laboratory [7], this shot included a secondary foil 9-cm downstream from the anode. Assuming that the virtual cathode would otherwise have formed at its usual position, 12-20 cm from the anode, the secondary foil would then have been between the anode and virtual cathode and in a position to extend the reflexing path length. The data for the other 13 shots with such an arrangement, with foil-to-foil spacings from 9 to 18 cm, do not show as clean or consistent spectra, however. Consequently, the effect of the second anode on the reflexing frequency is not clear in the data. In this experiment we also varied the anode/cathode gap from 16 through 23 cm, and we varied the anode/secondary-foil separation from 46 to 76 cm. At this spacing the secondary foil and anode would have formed a resonant cavity. The results have not demonstrated a strong microwave extraction dependence on foil positioning, which might be related to the severe overmoding in the source region.

A strong virtual cathode forms only when the diode current is much greater than the space-charge-limiting current  $I_{SCL} \simeq 17 \times (\gamma^{2/3} - 1)^{3/2}/G(kA)$ , where  $G \sim 1 - 2$  is a correction factor which might include two-dimensional and foil-scattering effects. Aurora's current of 300 kA only exceeded the space-charge limit by a factor of 2-4. The downstream current is an estimate of the spacecharge limit. We have estimated the charge deposition downstream by measuring the dose deposited in thermoluminescent detectors by the electron beam which penetrated the aluminum drift tube wall. These measurements [11], which suggest that the average current downstream was on the order of 60 kA, agree with previous CVR measurements and will be compared to ongoing PIC computations in future work.

# **IV.** CONCLUSIONS

We have demonstrated the extraction of ~50 J per shot of microwave radiation in the 0.6- to 1.0-GHz range into each of eight radial arms. The average energy per arm agrees reasonably with the observations of the previous Aurora microwave experiments [3]–[6]. With eight arms, we have the capability for radiating 400 J from singlemoded rectangular waveguides.

The original work on the Aurora reflex diode produced  $\sim 1$  kJ of microwaves as seen by an axial calorimeter, unfortunately obscured by a large background. To avoid this background problem we have investigated more recently the energy which could be extracted radially in rectangular waveguides. This approach has vielded reasonably good data on the energy per pulse, at least regarding in-band energy content. Up to a total of eight radial arms, the energy per radial arm has not diminished with an increasing number of arms, which suggests that we can extract more energy by simply adding arms. We are preparing for an experiment to extract energy into 18 radial arms, which might push the total energy extracted radially toward 1 kJ, in the absence of source saturation. With eight arms our extraction efficiency is much lower than experiments with axial extraction [12].

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